

Physics Informed Neural Networks (PINN) for PDE solving based on the simulation of a ray tracing method

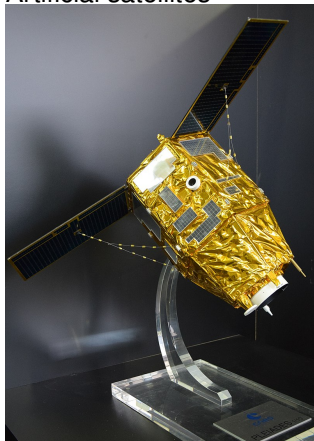
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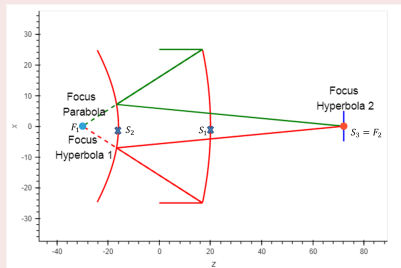
MSGI, November 26, 2021

Context

Artificial satellites



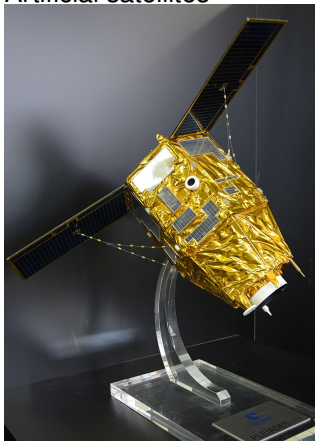
17th century: approach 1



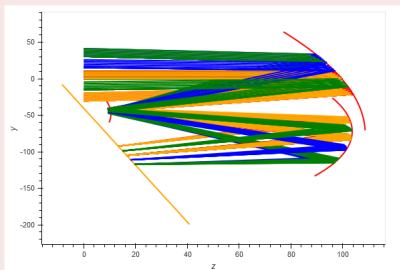
- ▶ Based on conic surfaces: parabola, hyperbola, ellipses
- ▶ Exact resolution
- ▶ Famous examples: Newton telescope (1 parabola)

Context

Artificial satellites



20th century: approach 2



- ▶ Use of freeform surfaces orthogonal polynomials or NURBS
- ▶ Parametric optimization problem
- ▶ Famous example: Three Mirrors Anastigmat (Gaia mission)

Problem

State-of-the-art

- ▶ Classical numerical methods such as Finite Elements
- ▶ Discrete approximation through statistical approach

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- ▶ Needs for alternative method
- ▶ Approximation of light intensity f
- ▶ Independant of discretization

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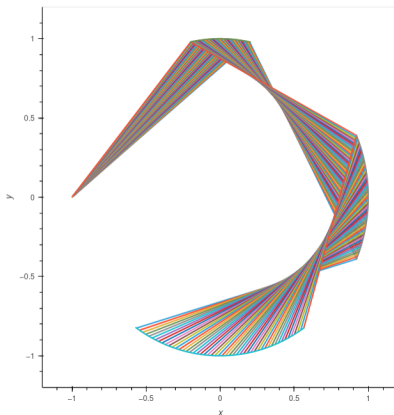
Work approaches

- ▶ Part 1: Statical approach
- ▶ Part 2: PINN approach

Part 1: Statistical methods

Modelling

Assuming that we know the light position p_0 and directions $\{d_1, \dots, d_N\}$ of N rays, the mirror locates at the unit circle around the origin, we evaluate the light intensity on the mirror $f(p, d_i), p \in \mathbb{R}^2 \cap \text{mirror}, d \in S^1$, using a statistical approach



Part 1: Statistical methods

The light intensity is define by:

$$f(p, d) = \lim_{N \rightarrow \infty} f^N(p, d)$$

With

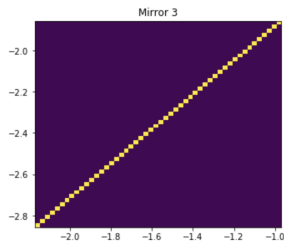
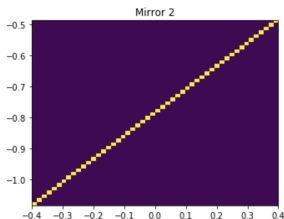
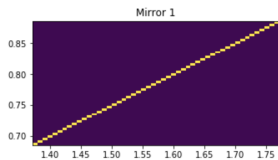
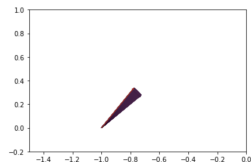
$$f^N(p, d) = \frac{1}{N} \sum_i \delta_{p_i, d_i}(p, d)$$

Question?

Make the histogram for the parameter pair (θ, d) , where θ is a parameterization of the mirror (S^1 in our case) and d the directions of the lights.

Single source on unit circle

- ▶ Source position: $p_0 = (-1, 0)$
- ▶ Direction:
 $\pi/4 - 0.1 < d < \pi/4 + 0.1$
- ▶ Number of rays: 1000

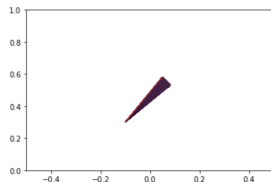


Remark 1

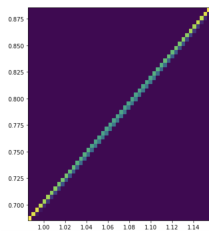
When the light source locates on the unit circle, the histogram is always uniform along the diagonals.

Single source inside unit circle

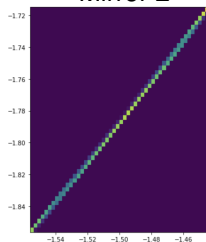
- ▶ Source position: $p_0 = (-0.1, 0.3)$
- ▶ Direction: $\pi/4 - 0.1 < d < \pi/4 + 0.1$
- ▶ Number of rays: 1000



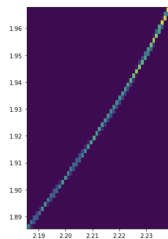
Mirror 1



Mirror 2



Mirror 3

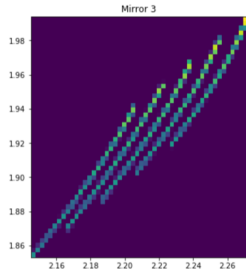
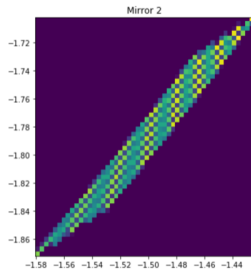
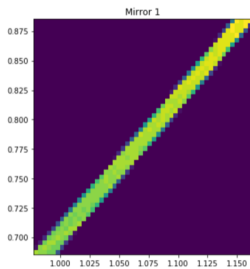
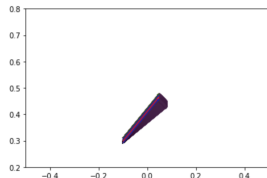


Remark 2

What happens if we have multiple sources?

Multiple sources on $x = -0.1$

- ▶ $p_0^y \in \{0.29, 0.295, 0.3, 0.305, 0.31\}$
- ▶ Direction: $\pi/4 - 0.1 < d < \pi/4 + 0.1$
- ▶ Number of rays: 5000

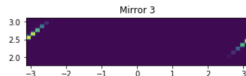
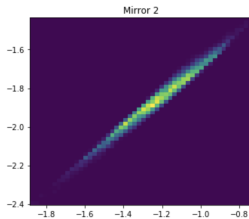
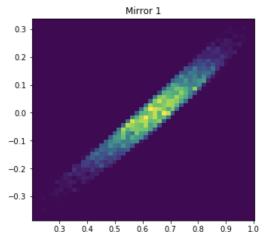
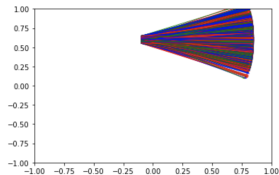


Remark 3

What happens if the source is no longer fixed?

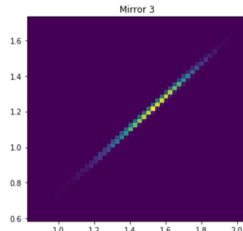
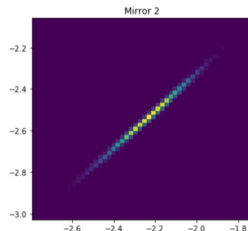
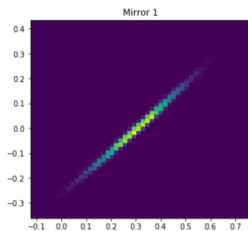
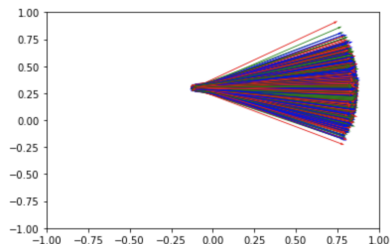
Uniform sources on $x = -0.1$

- ▶ $p_0^y = \text{Uniform}(0.29, 0.31)$
- ▶ Direction: $\pi/4 - 0.1 < d < \pi/4 + 0.1$
- ▶ Number of rays: 10000



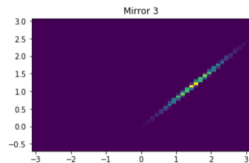
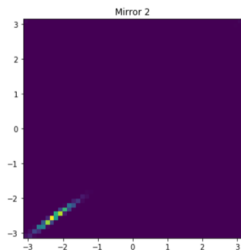
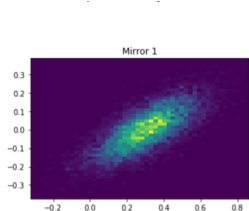
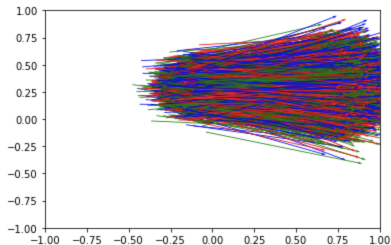
Gaussian source with small variance

- ▶ Source mean position: $(-0.1, 0.3)$
- ▶ Source variance matrix: $\begin{bmatrix} 0.0001 & 0 \\ 0 & 0.0001 \end{bmatrix}$
- ▶ Light mean: 0
- ▶ Light standard deviation: 0.1
- ▶ Number of rays: 10000



Gaussian source with large variance

- ▶ Source mean position: $(-0.1, 0.3)$
- ▶ Source variance matrix: $\begin{bmatrix} 0.01 & 0 \\ 0 & 0.01 \end{bmatrix}$
- ▶ Light mean: 0
- ▶ Light standard deviation: 0.1
- ▶ Number of rays: 10000



Part 2: PINN Approach

Modeling

We suppose that the equation is time independent and the density function f satisfies the following simplified version of a Liouville-Boltzmann equation:

$$\left\{ \begin{array}{l} d \cdot \nabla_p f + \lambda f = 0, \quad \text{on } \Omega, \quad 0 < \lambda \ll 1, \\ f(p, d) = f(p, d - 2(d, n)n), \quad \text{on } \Gamma_-^M, \\ f(p, d) = \mathbf{G}_{\sigma_p}(\|p - p_0\|) \times \mathbf{G}_{\sigma_d}(\|d - d_0\|), \quad \text{on } \Gamma_-^S. \end{array} \right.$$

where $(p, d) \in \Omega = \mathbf{D} \setminus \bar{W}_n \times \mathcal{S}^1$ and

$$\left\{ \begin{array}{l} \Gamma^+ = \{(p, d) \in \Omega, n(p) \cdot d \geq 0\}, \\ \Gamma^- = \{(p, d) \in \Omega, n(p) \cdot d < 0\}, \\ \Gamma_-^M = \{(p, d) \in \Gamma^-, p \in M\}, \\ \Gamma_-^S = \{(p, d) \in \Gamma^-, p \in S\}. \end{array} \right.$$

Motivation for PINN method

PINN advantages

- ▶ Once the learning step is completed, the resolution is direct, and requires no additional time
- ▶ The ability of parallel processing
- ▶ Cost and Time Benefits for enterprise

Physical problem

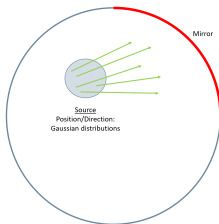


Figure: Schematic problem

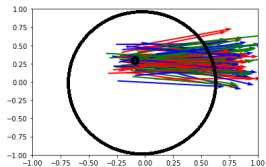


Figure: Rays shot

Question

Propose a PINN which makes it possible to approximate $f(p, d)$.

PINN Approach

General Loss formulation:

$$\mathcal{L} = \omega_{col}\mathcal{L}_{col} + \omega_{res}\mathcal{L}_{res} + \omega_{bc}\mathcal{L}_{bc} + \omega_{ic}\mathcal{L}_{ic}$$

Estimation

We propose a PINN estimating the density value $f(\theta, \varphi)$:

- ▶ PINN: MLP with 6 layers, TanH activation
- ▶ Loss: $\mathcal{L} = \omega_{col}\mathcal{L}_{col} + \omega_{bc}\mathcal{L}_{bc}$
- ▶ training set generated with statistical approach

Tests PINN

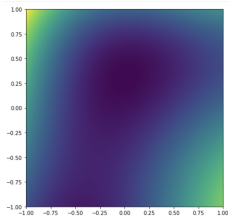


Figure: 1k epochs

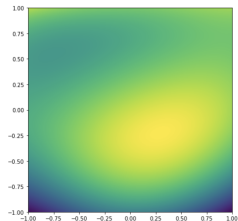


Figure: 2k epochs

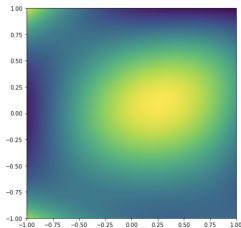


Figure: 20k epochs

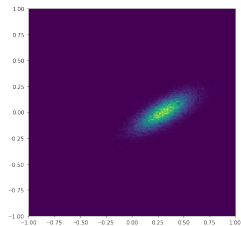


Figure: Groundtruth

Conclusion

Possible directions:

- ▶ Consider other activation functions and network structures
- ▶ If specific interest for surface values, possibility to take into account only those points
- ▶ For using residuals, consider the former equation and avoid trivial solution



**Thank
You!**